



An analytic hydrodynamical model of rotating 3D expansion in heavy-ion collisions

M. I. Nagy¹, T. Csörgő^{2,3},

¹Dept. of Atomic Physics, ELTE, H-1117 Budapest XI, Pázmány P. 1/A, Hungary

²Wigner RCP, H-1525 Budapest 114, POB 49, Hungary

³KRF, H-3200 Gyöngyös, Mátrai út 36, Hungary

December 4, 2015

Abstract

A new exact and analytic solution of non-relativistic fireball hydrodynamics is presented. It describes an expanding triaxial ellipsoid that rotates around one of its principal axes. The observables are calculated using simple analytic formulas. Azimuthal oscillation of the off-diagonal Bertsch-Pratt radii of Bose-Einstein correlations as well as rapidity dependent directed and third flow measurements provide means to determine the magnitude of the rotation of the fireball. Observing this rotation and its dependence on collision energy may lead to new information on the equation of state of the strongly interacting quark gluon plasma produced in high energy heavy ion collisions.

1 Introduction

The applications of hydrodynamics (both numerical and exact analytic solutions) in the description of the „soft” particle production in heavy-ion collisions has a good tradition. Here we deal with an analytic model. (For a recent work on numerical hydrodynamics, see eg. Ref. [1]; the many references therein provide a good list of recent developments in numerical hydrodynamical modelling). Exact solutions require some „luck” to find, however, they also have a good line of success: the early days of high energy physics saw the development of the Landau-Khalatnikov solution [2] and the Hwa-Bjorken solution [3], both invaluable useful for high energy collision phenomenology. Since then, many physically realistic solutions have been found, relativistic as well as non-relativistic ones.

The rotation of the expanding hot and dense matter produced in non-central heavy ion collisions drew much theoretical attention recently. Many numerical models try to grasp the effects of rotation on the observables (see eg. Refs. [4, 5]). In this work we present an exact solution of the non-relativistic perfect fluid hydrodynamical equations that describes the desired rotating expansion of a triaxial ellipsoid-like fireball, just as one imagines the space-time picture of a non-central heavy-ion collision. We also calculate the observables using simple analytic formulas and point out those features of them that would reveal the magnitude of rotation.

We mention some of our works of which this step is a natural continuation: the discovery of accelerating relativistic solutions [6, 7] as well as the first rotating relativistic [8] and non-relativistic [9] exact solutions. There are also exact solutions available for arbitrary equation of state in the relativistic [10] domain (even ones containing multipole asymmetries [11]) as well as in the non-relativistic domain [12] with realistic, ellipsoid-like symmetry. The solution and the calculation of the observables presented below is thus a next example of usefulness of exact solutions in heavy-ion phenomenology. The current work is a straightforward, although rather non-trivial generalization of our recent work on the evaluation of the observables for a spheroidally symmetric, rotating and expanding family of exact solutions of fireball hydrodynamics [13] for more general expansions with triaxial ellipsoidal symmetry. We refer to this work also for a more detailed overview and introduction to the status of the field of exact solutions of fireball hydrodynamics [13].

2 A new rotating solution of hydrodynamics

The equations of non-relativistic hydrodynamics are the Euler equation, the energy conservation equation and the particle number conservation equation. They are recited here in the form suitable for heavy-ion physics phenomenology:

$$nm_0 (\partial_t + \mathbf{v} \nabla) \mathbf{v} = -\nabla p, \quad (1)$$

$$(\partial_t + \mathbf{v} \nabla) \varepsilon = -(\varepsilon + p) \nabla \mathbf{v}, \quad (2)$$

$$(\partial_t + \mathbf{v} \nabla) n = -n \nabla \mathbf{v}. \quad (3)$$

Here n stands for the particle number density (thus nm_0 is the mass density), T for temperature, p for pressure, \mathbf{v} for the velocity field, ε for the energy density. The equations above need to be supplemented with a suitable Equation of State (EoS) which we customarily choose as

$$\varepsilon = \kappa(T) p, \quad (4)$$

$$p = nT. \quad (5)$$

In non-central heavy-ion collisions, an almond-shaped region of hot and dense matter forms which has non-zero angular momentum. The almond shape is approximated with a triaxial ellipsoid, and the following solution of hydrodynamics takes rotation and expansion into account. Let the x axis point in the direction of the impact parameter and the z axis in the direction of the colliding beams; thus the x - z plane is the event plane, and the rotation is around the y axis. We call the lab frame the K frame, to be distinguished from the K' frame, which co-rotates with the principal axes of the expanding ellipsoid. The solution is described in terms of the time-dependent angle of rotation $\vartheta(t)$, and the three principal axes of the rotated ellipsoid $X(t)$, $Y(t)$, $Z(t)$. The scaling variable, whose level surfaces correspond to self-similar ellipsoids is¹

$$s(\mathbf{r}, t) \equiv \frac{r'^2_x}{X^2} + \frac{r'^2_y}{Y^2} + \frac{r'^2_z}{Z^2}, \quad (6)$$

¹In the following, we do not explicitly write up the time dependence of functions which were introduced as time-dependent ones.

$$\begin{aligned}
r'_x &= r_x \cos \vartheta - r_z \sin \vartheta, \\
r'_y &= r_y, \\
r'_z &= r_x \sin \vartheta + r_z \cos \vartheta.
\end{aligned} \tag{7}$$

The velocity field is taken to be a linear, rotating Hubble-like one, for which the above s variable is a proper „scaling variable”, ie. its co-moving derivative vanishes. We introduce the $V = XYZ$ notation. One can directly check that for constant κ the above and the following formulas provide the desired rotating solution to the (1)–(3) hydrodynamical equations:

$$\mathbf{v}'(\mathbf{r}, t) = \begin{pmatrix} \frac{\dot{X}}{X} r'_x + g \frac{X}{Z} r'_z \\ \frac{\dot{Y}}{Y} r'_y \\ \frac{\dot{Z}}{Z} r'_z - g \frac{Z}{X} r'_x \end{pmatrix}, \tag{8}$$

$$\dot{\vartheta}(t) = g(t) = \frac{\omega_0}{2} \left(\frac{X_0 + Z_0}{X + Z} \right)^2, \tag{9}$$

$$n = n_0 \frac{V_0}{V} \mathcal{V}(s), \tag{10}$$

$$T = T_0 \left(\frac{V_0}{V} \right)^{\frac{1}{\kappa}} \mathcal{T}(s), \tag{11}$$

$$\mathcal{V}(s) = \frac{1}{\mathcal{T}(s)} e^{-\frac{1}{2} \int_0^s \frac{ds'}{\mathcal{T}(s')}}. \tag{12}$$

$$\tag{13}$$

For non-constant κ , the above formulas can be modified to form a valid solution. A solution is found with spatially homogeneous T and Gaussian-like n as

$$n = n_0 \frac{V_0}{V} e^{-s/2}, \tag{14}$$

$$T \equiv T(t), \quad \frac{V_0}{V} = \exp \left\{ \int_{T_0}^T d\beta \left(\frac{d\kappa}{d\beta} + \frac{\kappa}{\beta} \right) \right\}. \tag{15}$$

Furthermore, the time development of the X, Y, Z principal axes (and thus that of $\dot{\vartheta}$ and g is governed by a Lagrangian for a point-like particle of mass m_0 , which we write up only in the $\kappa = \text{const}$, $T \equiv T(t)$ case:

$$L = \frac{m_0}{2} (\dot{X}^2 + \dot{Y}^2 + \dot{Z}^2) - \frac{1}{\kappa} \left(\frac{T}{T_0} \right) - m_0 \omega^2 R_0^2, \tag{16}$$

where we introduced another simplifying notation: the „average” radius of the ellipsoid R , and the „average” angular velocity $\omega(t)$ as

$$R \equiv \frac{X + Z}{2}, \quad R_0 \equiv \frac{X_0 + Z_0}{2}, \quad \omega(t) \equiv \omega_0 \frac{R_0^2}{R^2}. \tag{17}$$

Note that the velocity profile is written up in the co-rotating K' frame. The free parameter ω_0 is related to the total angular momentum M_y of the flow (which points in the y direction); for example, in the $\mathcal{V}(s) = e^{-s/2}$ Gaussian case

$$M_y = 2M_0 R_0^2 \omega_0, \tag{18}$$

$$M_0 = m_0 (2\pi)^{3/2} V_0 n_0. \tag{19}$$

The solution thus specified is a natural generalization of the earlier results for three dimensionally expanding, non-rotating triaxial ellipsoids [12, 14] (indeed, our formulation resembles very much to those ones), and in the spheroidal limit ($X = Z$) it reproduces the exact solutions presented in Refs. [9, 13] for rotating and exploding fireballs with spheroidal symmetry.

3 Observables from the new solution

One can calculate the hadronic observables from the above solution by specifying the emission function (source function) and a suitable freeze-out condition. Following Ref. [12], for simplicity (and because the solution for arbitrary $\kappa(T)$ is then available), we take a Gaussian ansatz for the density profile (thus the temperature is spatially homogeneous), as well as take the freeze-out to happen at a constant time t_f (which in this case corresponds to a given T_f freeze-out temperature). We assume that at the freeze-out, particles with mass m are created. The source function is then taken to be the non-relativistic Boltzmann distribution for a particle with mass m :

$$S(\mathbf{r}, \mathbf{p}) = \frac{n(t_f, \mathbf{r})}{(2\pi m T_f)^{3/2}} \exp \left\{ -\frac{(\mathbf{p} - m\mathbf{v}(t_f, \mathbf{r}))^2}{2mT(t_f)} \right\}. \quad (20)$$

In the following, the $_f$ index means the value taken at the freeze-out time t_f , but it is mostly dropped: all quantities are to be understood as their value at t_f . The single-particle spectrum and the two-particle Bose-Einstein correlation function are calculated as

$$N_1(\mathbf{p}) = \int d^3\mathbf{r} S(\mathbf{r}, \mathbf{p}), \quad (21)$$

$$C(\mathbf{K}, \mathbf{q}) = 1 + \lambda \frac{|\tilde{S}(\mathbf{K}, \mathbf{q})|^2}{|\tilde{S}(\mathbf{K}, \mathbf{0})|^2}, \quad (22)$$

where \mathbf{K} and \mathbf{q} are the average pair momentum and the relative momentum, respectively, and λ is the effective intercept parameter of the correlation function. For the mentioned Gaussian density case, all the integrals can be performed analytically, yielding simple results. We introduce the following quantities for convenience:

$$T_{xx}^* = T_f + m(\dot{X}^2 + \omega^2 R^2), \quad (23)$$

$$T_{zz}^* = T_f + m(\dot{Z}^2 + \omega^2 R^2), \quad (24)$$

$$T_{xz}^* = m\omega R(\dot{X} - \dot{Z}). \quad (25)$$

Note that these definitions can be straightforwardly specialized to the case when the two principal axes X and Z are equal ($X = Z = R$), and thus one deals with the expansion of a rotating spheroid.

The single particle spectrum can be then written as

$$\begin{aligned} \frac{dn}{d^3\mathbf{p}} &\propto \exp \left(-\frac{p_x^2}{2mT_x} - \frac{p_y^2}{2mT_y} - \frac{p_z^2}{2mT_z} - \frac{\beta_{xz}}{m} p_x p_z \right) = \\ &= \exp \left(-\frac{p_x'^2}{2mT_x'} - \frac{p_y'^2}{2mT_y'} - \frac{p_z'^2}{2mT_z'} - \frac{\beta'_{xz}}{m} p_x' p_z' \right), \end{aligned} \quad (26)$$

where T_x , T_y , T_z , and β_{xz} , as well as T_x' , T_y' , T_z' and β'_{xz} characterize the slope parameters (and the cross-terms) in the K (laboratory) frame and the K' frame (the eigen-frame of the rotated ellipsoid), respectively. The expression of these parameters are given as

$$T_x' = T_{xx}^* - \frac{T_{xz}^{*2}}{T_{zz}^*}, \quad (27)$$

$$T_z' = T_{zz}^* - \frac{T_{xz}^{*2}}{T_{xx}^*}, \quad (28)$$

$$\beta'_{xz} = \frac{T_{xz}^*}{T_{xx}^* T_{zz}^* - T_{xz}^{*2}}, \quad (29)$$

$$\frac{1}{T_x} = \frac{\cos^2 \vartheta}{T'_x} + \frac{\sin^2 \vartheta}{T'_z} + \beta'_{xz} \sin(2\vartheta), \quad (30)$$

$$\frac{1}{T_z} = \frac{\sin^2 \vartheta}{T'_x} + \frac{\cos^2 \vartheta}{T'_z} - \beta'_{xz} \sin(2\vartheta), \quad (31)$$

$$\beta_{xz} = \frac{\sin(2\vartheta)}{2} \left(\frac{1}{T'_z} - \frac{1}{T'_x} \right) + \beta'_{xz} \cos(2\vartheta), \quad (32)$$

$$T'_y = T_y = T + m\dot{Y}^2. \quad (33)$$

We see that the formulas relating the parameters in the K and K' frames are simply the ones describing a rotation by a fixed θ angle (the value of the tilt angle at freeze-out). The fact that a cross term (ie. nonzero β'_{xz}) appears (this is a new feature compared to Ref. [12]) signals that the single-particle spectrum is *not* diagonal in the K' frame: the tilt of the coordinate-space ellipsoid is not the same as that of the ellipsoid determined from the single-particle spectrum.

The expression of the azimuthal angle-averaged p_T spectrum and the anisotropy (v_n) parameters follow the footsteps of Ref. [12]: one can introduce the scaling variables v and w and the effective temperature T_{eff} as

$$v \equiv -\frac{\beta_{xz}}{m} p_z p_T, \quad w \equiv \frac{p_T^2}{4m} \left(\frac{1}{T_y} - \frac{1}{T_x} \right), \quad \frac{1}{T_{\text{eff}}} = \frac{1}{2} \left(\frac{1}{T_y} + \frac{1}{T_x} \right). \quad (34)$$

Using the $I_n \equiv I_n(w) = \frac{1}{\pi} \int_0^\pi \cos(n\varphi) e^{w \cos \varphi} d\varphi$ modified Bessel functions of w , up to second order in v (and thus in the y rapidity) one has

$$E \frac{dn}{d^3\mathbf{p}} = \frac{1}{2\pi p_T} \frac{dn}{dp_T dy} \left\{ 1 + 2 \sum_{n=1}^{\infty} v_n \cos[n(\varphi - \Psi_n)] \right\} \Rightarrow, \quad (35)$$

$$\frac{1}{2\pi p_T} \frac{dn}{dp_T dy} \propto \exp \left\{ -\frac{p_z^2}{2mT_z} - \frac{p_T^2}{2mT_{\text{eff}}} \right\} \times \left[I_0(w) + \frac{I_0(w) + I_1(w)}{4} v^2 \right], \quad (36)$$

$$v_1 = \frac{v}{2} \frac{I_0(w) + I_1(w)}{I_0(w)}, \quad (37)$$

$$v_2 = \frac{I_1(w)}{I_0(w)} + \frac{v^2}{8} \left[1 + \frac{I_2(w)}{I_0(w)} - 2 \frac{I_1^2(w)}{I_0^2(w)} \right], \quad (38)$$

$$v_3 = \frac{v}{2} \frac{I_1(w) + I_2(w)}{I_0(w)}. \quad (39)$$

Here φ denotes the azimuthal angle, Ψ_n is the n th order event plane (in our model, all the event planes coincide, ie. fluctuation effects are not taken into account). Fig. 1 presents the characteristic rapidity dependence of v_1 , v_2 and v_3 . The two-particle Bose-Einstein correlation also turns out to be a Gaussian, and can either be expressed in terms of the \mathbf{q}' of the \mathbf{q} relative momentum (in the K and K' frames, respectively), or using the Bertsch-Pratt parametrization. Simple calculation yields

$$C(\mathbf{K}, \mathbf{q}) = 1 + \lambda \exp \left\{ - \sum_{ij} q'_i q'_j R'^2_{ij} \right\} = 1 + \lambda \exp \left\{ - \sum_{ij} q_i q_j R^2_{ij} \right\}, \quad (40)$$

$$R'^2_x = X^2 \frac{T_f}{T'_x}, \quad (41)$$

$$R'^2_z = Z^2 \frac{T_f}{T'_z}, \quad (42)$$

$$R'^2_{xz} = XZ T_f \beta'_{xz}, \quad (43)$$

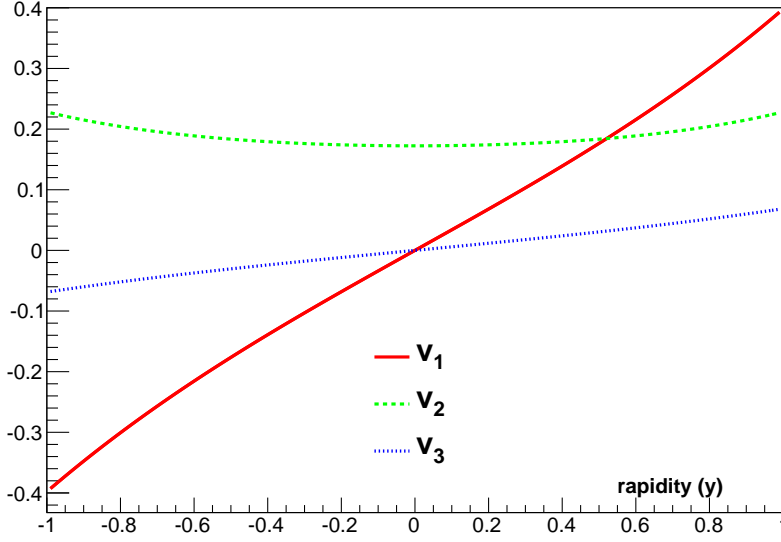


Figure 1: Dependence of the azimuthal anisotropy parameters v_1 , v_2 and v_3 on rapidity y , for a reasonable choice of the freeze-out parameters (namely: $\vartheta_f = \pi/5$, $\beta'_{xz} = 5 \cdot 10^{-4}$ 1/MeV, $p_T = 300$ MeV/c, $m = 140$ MeV, $T_x = 270$ MeV, $T_y = 170$ MeV, $T_z = 600$ MeV). The dependence of v_1 and v_3 is characteristic to a system tilted by some ϑ_f angle.

where we utilized the already calculated values of the T'_x , T'_z , β'_{xz} slope parameters. These radii are valid in the K' frame; transformation into the K frame (lab frame) is straightforward:

$$R_x^2 = R_x'^2 \cos^2 \vartheta + R_z'^2 \sin^2 \vartheta + R_{xz}'^2 \sin(2\vartheta), \quad (44)$$

$$R_z^2 = R_x'^2 \sin^2 \vartheta + R_z'^2 \cos^2 \vartheta - R_{xz}'^2 \sin(2\vartheta), \quad (45)$$

$$R_{xz}^2 = \frac{\sin(2\vartheta)}{2} (R_x'^2 - R_z'^2) + \cos(2\vartheta) R_{xz}'^2. \quad (46)$$

The y direction, as before, is much simpler:

$$R_y'^2 = R_y^2 = Y^2 \frac{T_f}{T_y}. \quad (47)$$

Customarily introducing the o , s , l (out, side, long) Bertsch-Pratt components of \mathbf{q} in the longitudinally co-moving system (LCMS), one has

$$C(\mathbf{K}, \mathbf{q}) = 1 + \lambda \exp \left\{ - \sum_{a,b=o,s,l} q_a q_b R_{ab}^2 \right\}, \quad (48)$$

and the radii as functions of the pair momentum azimuthal angle φ are

$$\begin{aligned} R_{oo}'^2 &\equiv R_o^2 = R_x^2 \cos^2 \varphi + R_y^2 \sin^2 \varphi, \\ R_{ss}'^2 &\equiv R_s^2 = R_x^2 \sin^2 \varphi + R_y^2 \cos^2 \varphi, \\ R_{ll}'^2 &\equiv R_l^2 = R_z^2, \\ R_{os}'^2 &= (R_y^2 - R_x^2) \sin \varphi \cos \varphi, \\ R_{ol}'^2 &= R_{xz}^2 \cos \varphi, \\ R_{sl}'^2 &= R_{xz}^2 \sin \varphi. \end{aligned} \quad (49)$$

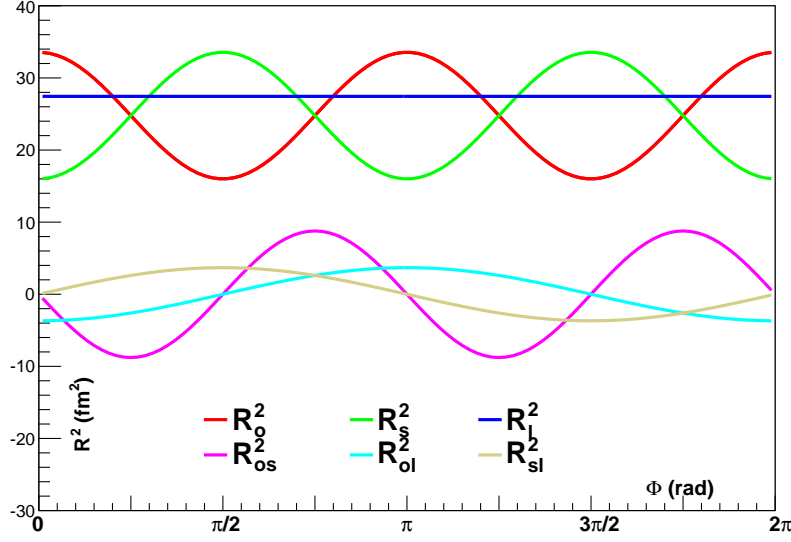


Figure 2: φ dependence of the Bertsch-Pratt radii for a reasonable choice of the freeze-out parameters (namely: $\vartheta_f = \pi/5$, $R_x'^2 = 25 \text{ fm}^2$, $R_y'^2 = 16 \text{ fm}^2$, $R_z'^2 = 36 \text{ fm}^2$, $R_{xz}'^2 = 5 \text{ fm}^2$). The oscillation of R_s , R_o and R_{os} with a π period is characteristic to an ellipsoid-shaped source. The 2π period oscillation of R_{ol} and R_{sl} are characteristic to a tilted (and rotating) source.

Fig. 2 shows the Bertsch-Pratt radii as a function of pair momentum azimuthal angle. The 2nd order oscillating ($\cos^2 \varphi$, $\sin^2 \varphi$ containing) BP-radii, also the cross-terms, were recently extensively measured by the STAR collaboration [15], however, as we see, it is not these oscillations that signal rotational expansion.

4 Discussion and outlook

Most of the formulae of the observables derived here show a very close similarity (even identity) to those calculated in Ref. [12], where a *stationary, but tilted* ellipsoidal source was assumed. The universal scaling of the elliptic flow and other anisotropies (a feature of experimental data that was successfully explained in terms of the Buda-Lund model [16]) is preserved here, just the scaling variables are expressed in a bit more involved way, see Eq. (34). The oscillation of the HBT radii characteristic to rotation also shows up in other models of tilted ellipsoidal sources (just as in Ref. [12]; a simpler model was introduced earlier in Ref. [17]).

The new result in the present work is two-fold: first, an actual (and in some sense, unique) rotating and expanding hydrodynamical solution is found which naturally leads to tilted sources from a non-tilted initial condition (and it can be applied to follow the time-evolution of rotation, for any given $\kappa(T)$ EoS). Second, the result for the observables are refined in a way that takes not only the tilt but also the rotational flow into account.

The most striking consequence of tilted (and also rotating) expansion are seen in the rapidity dependence of the anisotropy parameters (Fig. 1) as well as the *directed flow-like* (ie. azimuthally 2π -periodic) oscillations of the R_{sl} and R_{ol} HBT radius parameters (the appearance of these cross-terms is purely due to rotation, see Eq. (49)). A natural proposal is thus to measure these observables to infer the rotational angle ϑ_f and possibly the angular velocity. It is worthwhile to mention that to do these measurements, one needs experimental information on not only the usual (2nd order) reaction plane

but also on the first order reaction plane; this may be done precisely by utilizing the y -dependence of the v_1 anisotropy parameter. It is also worth to mention that in the case of models with spheroidal expansion (that was discussed eg. in Refs. [4, 9]), the „angle of rotation” ϑ_f becomes ill-defined, so the experimental signatures of rotation become much harder to be identified.

The detection of rotation has some far-going promises. A softer equation of state (caused eg. by the presence of a critical endpoint on the QCD phase diagram) would mean that the matter expands less violently, thus the angle of rotation will be greater. In this manner, measuring the rotation angle as a function of collision energy can be of great importance in mapping out the critical endpoint and the location of quark-hadron phase transition. So we look forward to see whether our presented model is applicable to new collision-energy dependent measurements of HBT radii oscillations as well as rapidity dependent anisotropy parameters to infer the rotation of the expanding hot and dense matter. It would lead to new knowledge of the strongly interacting quark-gluon-plasma produced in heavy-ion collisions.

This work was supported by the Hungarian OTKA grant NK101438. M. N. was supported by the TÁMOP 4.2.4. A/1-11-1-2012-0001 „National Excellence Program”, financed by the European Union and the State of Hungary, co-financed by the European Social Fund.

References

- [1] I. Karpenko, P. Huovinen and M. Bleicher, Comput. Phys. Commun. **185**, 3016 (2014)
- [2] L. D. Landau, Izv. Akad. Nauk Ser. Fiz. **17**, 51 (1953),
I.M. Khalatnikov, Zhur. Eksp. Teor. Fiz. **27**, 529 (1954),
S. Z. Belenkij, L. D. Landau, Nuovo Cim. Suppl. **3S10**, 15 (1956).
- [3] R. C. Hwa, Phys. Rev. D **10**, 2260 (1974),
J. D. Bjorken, Phys. Rev. D **27**, 140 (1983).
- [4] L. P. Csernai, S. Velle and D. J. Wang, Phys. Rev. C **89**, 034916 (2014)
- [5] L. P. Csernai, D. J. Wang and T. Csörgő, Phys. Rev. C **90**, 024901 (2014)
- [6] T. Csörgő, M. I. Nagy and M. Csanád, Phys. Lett. B **663**, 306 (2008)
- [7] M. I. Nagy, T. Csörgő and M. Csanád, Phys. Rev. C **77**, 024908 (2008)
- [8] M. I. Nagy, Phys. Rev. C **83**, 054901 (2011)
- [9] T. Csörgő and M. I. Nagy, Phys. Rev. C **89**, 044901 (2014)
- [10] M. Csanád, M. I. Nagy and S. Lökös, Eur. Phys. J. A **48**, 173 (2012)
- [11] M. Csanád and A. Szabó, Phys. Rev. C **90**, no. 5, 054911 (2014)
- [12] T. Csörgő, S. V. Akkelin, Y. Hama, B. Lukács and Y. M. Sinyukov, Phys. Rev. C **67**, 034904 (2003)
- [13] T. Csörgő, M. I. Nagy and I. F. Barna, arXiv:1511.02593
- [14] T. Csörgő, Acta Phys. Polon. B **37**, 483 (2006)
- [15] L. Adamczyk *et al.* [STAR Collaboration], Phys. Rev. C **92**, 014904 (2015)
- [16] M. Csanád, B. Tomasik and T. Csörgő, Eur. Phys. J. A **37**, 111 (2008)
- [17] M. A. Lisa, U. W. Heinz and U. A. Wiedemann, Phys. Lett. B **489**, 287 (2000)